



Deep Fields Planning



- At this meeting:
 - each breakout group would aim to identify key science drivers, and map these to desired observational capabilities
 - summary session would capture synergies and differences
- After this meeting (near-term):
 - white papers laying out in more detail the science drivers and the necessary observational capabilities
 - in context of LSST drilling fields call, and also other facilities
- After this meeting (longer term):
 - sub-groups, centered on various science goals, continue discussions in context of all the observations/ facilities needed



High-level observational needs



- Mapping the science drivers to desired observations – need to quantify depth, bandpasses, resolution, cadence:
 - imaging
 - spectroscopic
 - temporal
- For all the observations, what do they need in terms of:
 - future facilities
 - current facilities, but future observations necessary
 - current observations already in hand
- Would the science benefit from 'joint processing' (currently envisioned for WFIRST/LSST/Euclid), if so how?



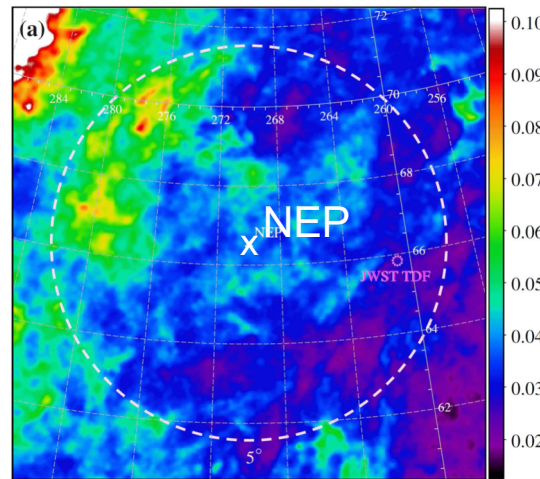
Spatial overlap, cadence



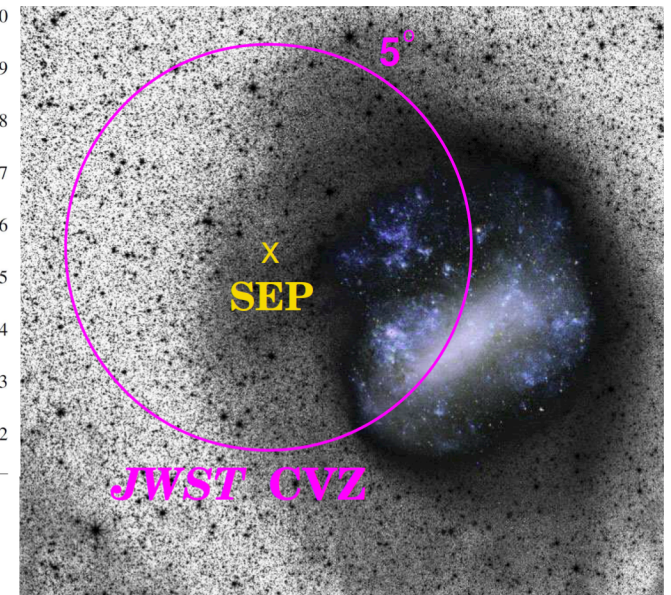
- If multiple facilities are needed for a given science goal, need to quantify overlap considerations in terms of:
 - sky coverage needed
 - temporal overlap needed
 - is CVZ required? (for WFIRST and/ or also Euclid, JWST or other facilities)
- Need to quantify whether currently existing fields can be used:
 - if so, which fields
 - which existing (or future) datasets on these fields would be most important?

Observatory Parameters

- CVZ for WFIRST, JWST, Euclid centered on ecliptic poles:
 - NEP: 18h 00m 00s +66° 33' 38.5520"
 - SEP: 06h 00m 00s -66° 33' 38.5520"
- CVZ radius is governed by sun angle constraints / field of regard, which can be different for each observatory:
 - WFIRST: $r = 36^\circ$
 - JWST: $r = 5^\circ$
 - Euclid: $r = 5^\circ$
- Note however that SEP is near LMC.



Jansen & Windhorst (2018);
Besla et al (2016).





Nominal Depth Capabilities



(see also talks by Bianco, Foley, Capak; and Hounsell et al 2017); sensitivities in all cases are total-depth 5σ point source AB mag:

• WFIRST (PSF $\sim 0.1''$):	Y	J	H	F184
HLS ($\sim 2,200 \text{ deg}^2$)	26.7	26.9	26.7	26.2
"SN medium" ($\sim 20 \text{ deg}^2$)	27.3*	27.7	27.6	27.0*
"SN deep" ($\sim 1\text{-}2 \text{ deg}^2$)	28.6*	29.0	28.2	27.7*

*These are not in the SN survey, but are representative, for similar exposure times

• Euclid (PSF $\sim 0.3''$):	Y	J	H
Wide ($\sim 15,000 \text{ deg}^2$)	24	24	24
Deep ($\sim 40 \text{ deg}^2$)	26	26	26

- LSST (PSF $\sim 0.7''$) Deep Drilling Fields (also Brandt 2016):
u=28.5, g=28.7, r=28.9, i=28.4, z=28.0, y=27.0



WFIRST CVZ, LSST and E(B-V)



- Map of $E(B-V) < 0.05$ in WFIRST CVZ
- green is $E(B-V) \sim 0.05$, white is $E(B-V) \sim 0$
- Including LSST coverage would exclude regions $> \text{dec } +40$

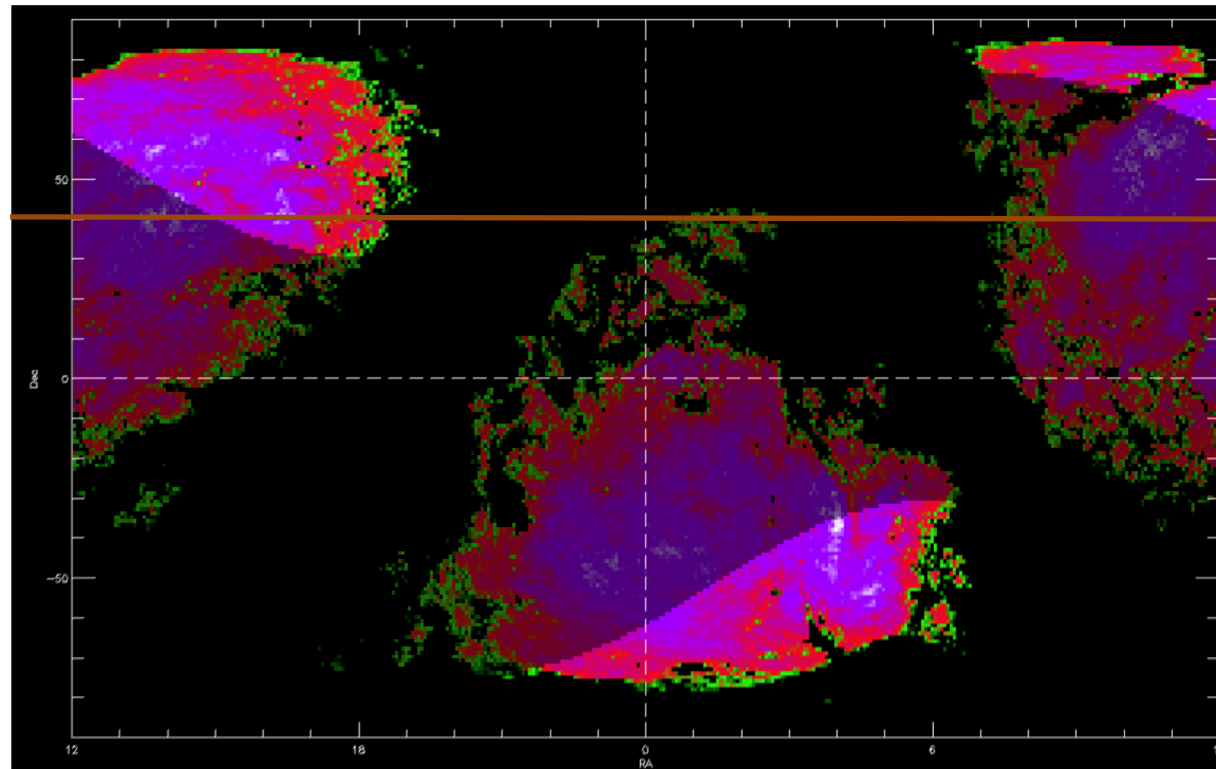


Figure courtesy of Ryan Foley



Sky coverage and current fields



- Blue lines indicate current CVZ limits for WFIRST, centered on ecliptic poles (based on sun angle constraint of 54°)

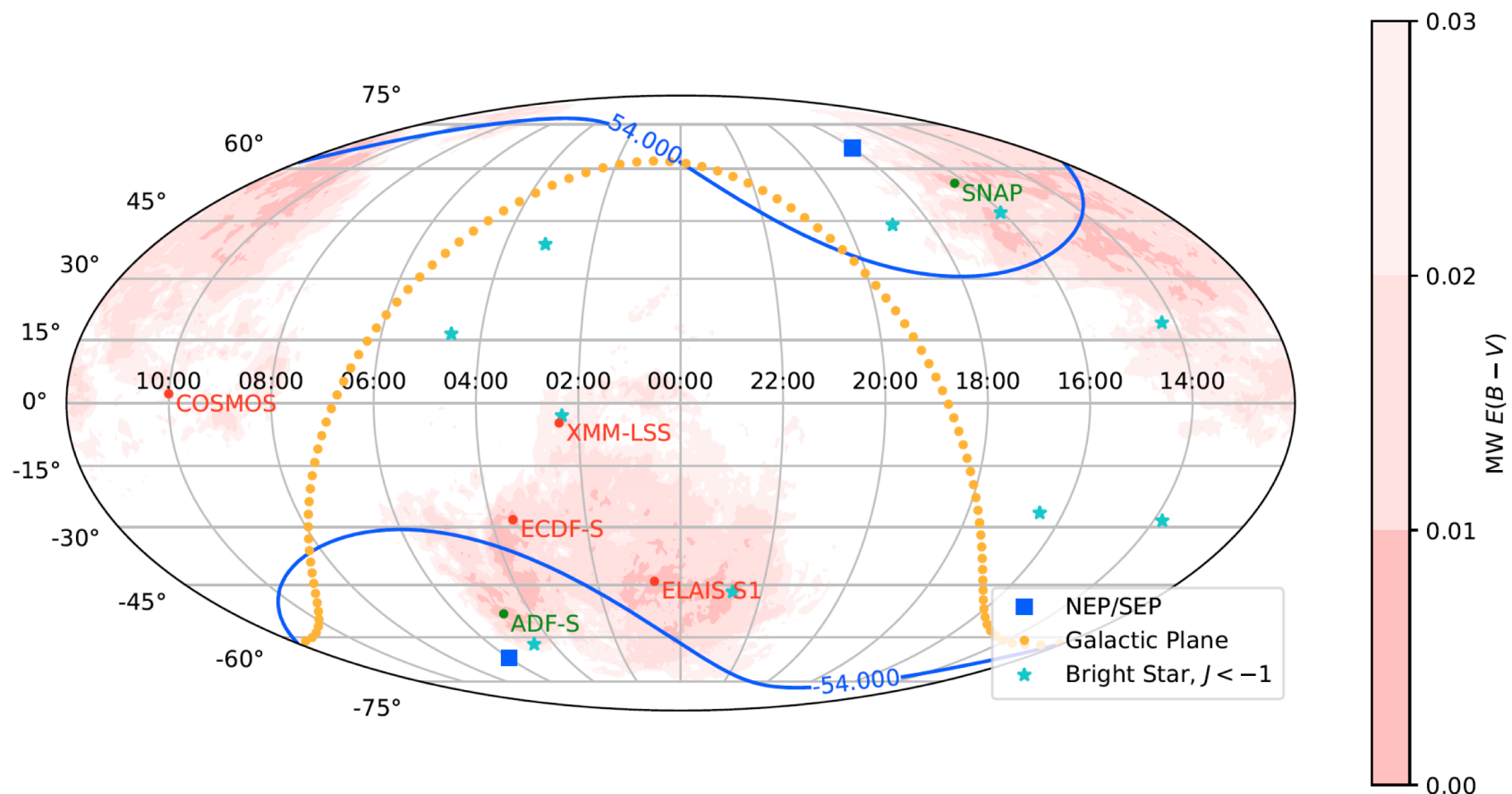


Figure courtesy of David Rubin



List of some current fields



<https://www.overleaf.com/10469108xwhqkxtcfycm#/39017211/>

Field	R.A.	Dec.	Ecl. Lat.	Area	E(B-V)	Rel. Zodi ^a	Days/Year
WFIRST CVZ fields (< 36°)							
IRAC Dark Field	17:40	+69:00	+87	0.2	0.043	1.0	365
Extended Groth Strip	14:17	+52:30	+60	0.2	0.009	1.2	365
GOODS-N	12:36	+62:13	+57	0.25	0.012	1.2	365
Deep2A	16:52	+34:55	+57	1	0.018	1.2	365
Elias N-2	16:46	+41:01	+63	5	0.014	1.1	365
Elias N-1	16:11	+55:00	+73	9	0.008	1.0	365
Akari Deep Field South	04:44	-52:20	-73	12	0.008	1.0	365
NEP-JWST-GTO-TDS	17:22	+65:49	+86	0.2	0.042	1.0	365
NEP-Spitzer	18:00	+66:33	+90	10	0.046	1.0	365
SEP-Spitzer	06:00	-66:33	-90	10	0.062	1.0	365
Non-CVZ fields							
CDFS	03:32	-27:48	-45	0.3	0.008	1.4	229
Deep2B	23:30	+00:00	+3	1	0.044	19	146
SSA22	22:17	+00:24	+10	4	0.056	5.6	149
COSMOS	10:00	+02:12	-9	2	0.018	6.0	148
VVDS14h	14:00	+05:00	+16	4	0.026	3.6	153
Elias S-1	00:35	-43:40	-43	7	0.008	1.5	215
Bootes	14:32	+34:16	+46	9	0.016	1.4	236
Lockman Hole	10:45	+58:00	+45	11	0.011	1.4	229
XMM-LSS	02:31	-04:30	-18	11	0.024	3.2	155
SPT Deep	23:30	-55:00	+46	100	0.010	1.4	236
HERA	07:00	-30:43		1200			

^aZodi values provided by David Rubin based on his calculations, corresponding to plane parallel zodi, relative to a normalized value of 1.0 at the ecliptic poles.