



The Van Cittert-Zernike Theorem: Theory, History, and Applications

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Grad Seminar (AST 542)

This *was* a talk on Radiative Transfer codes...

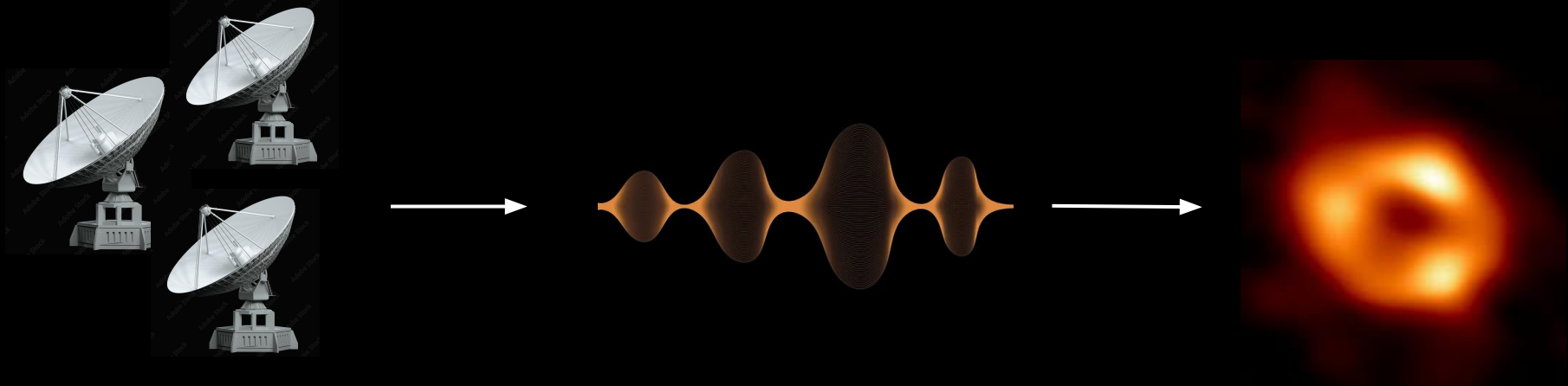
(but Nick and Philippe did such good jobs that I had to pivot.)

- **Background:** basics of interferometry and coherence theory.
- **Mathematical Foundations:** the van Cittert-Zernike Theorem.
- **Broader Applications:** the van Cittert-Zernike Theorem for an example interferometric array.
- **Resources:** simple coding example and how to find interferometric data.

Warning: includes a lot (and I mean *a lot*) of clickable links.

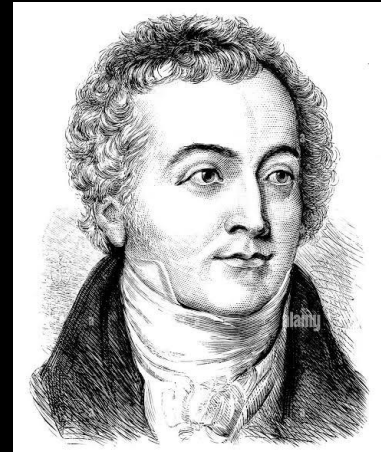
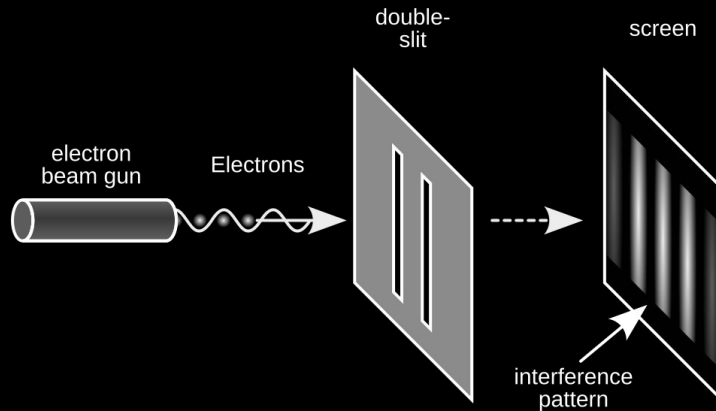
The golden rule: multiple telescopes are stronger than one.

- Simply, interferometry works by combining two or more signals to create a pattern to extract information about objects.



Coherence and interference...

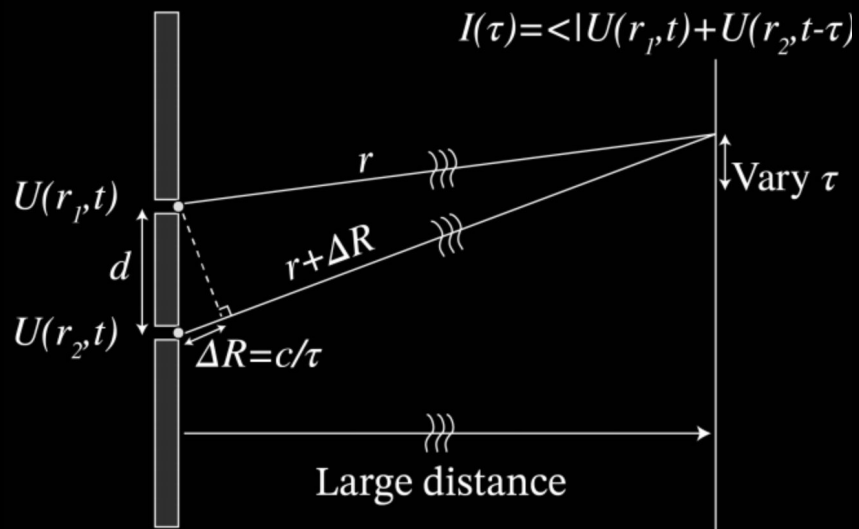
- **Young's Double-Slit Experiment:** interference pattern shows that light is a wave.
 - Temporal and Spatial.



A very serious portrait of
Thomas Young

[More on Young's Double Slit Experiment](#)

Mutual Spatial Coherence Function



Mutual Spatial Coherence Function

- Understanding the spatial coherence of two points.

$$\tau = \frac{\Delta R}{c}.$$

$$\begin{aligned} I(\tau) &= \langle |U_1(\mathbf{r}, t) + U_2(\mathbf{r}, t)|^2 \rangle \\ &= \langle |U(\mathbf{r}_1, t)|^2 \rangle + \langle |U(\mathbf{r}_2, t - \tau)|^2 \rangle + 2 \operatorname{Re} \langle U(\mathbf{r}_1, t) U(\mathbf{r}_2, t - \tau)^* \rangle. \end{aligned}$$

$$\Gamma_{12}(\tau) = \langle U(\mathbf{r}_1, t) U(\mathbf{r}_2, t - \tau)^* \rangle,$$

The Van Cittert-Zernike Theorem

- The “heart” of interferometry.
- **Spoiler alert:** there exists a relationship between the mutual spatial coherence function and the sky intensity distribution.
- Ducks?



Very serious portraits of
Pieter Hendrik van Cittert and Frits Zernike

The Van Cittert-Zernike Theorem



The Van Cittert-Zernike Theorem

$$\mathcal{V}(u,v) = \iint I(l,m) e^{-2\pi i(ul + vm)} dl dm$$

Mutual spatial coherence function
or the Visibility function

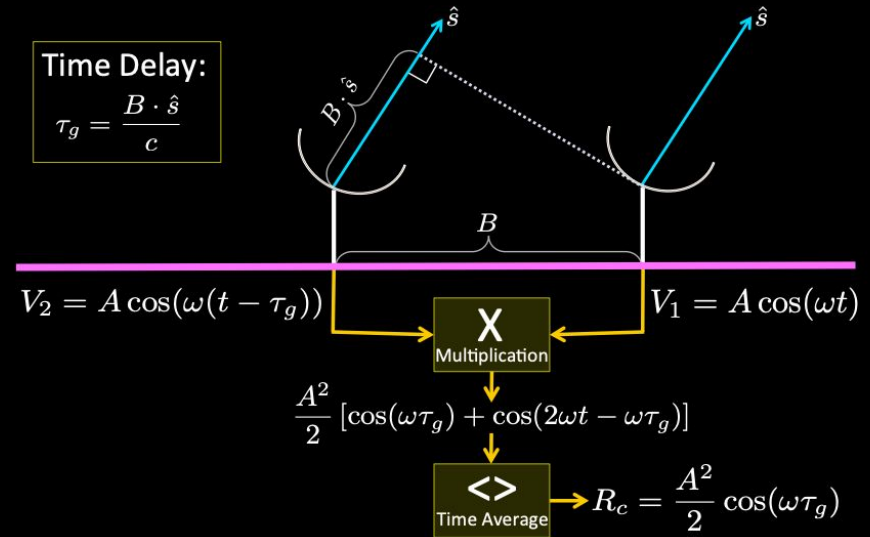
Sky intensity
distribution



A less serious portrait of
Joseph Fourier

Van Cittert-Zernike Theorem in Interferometry

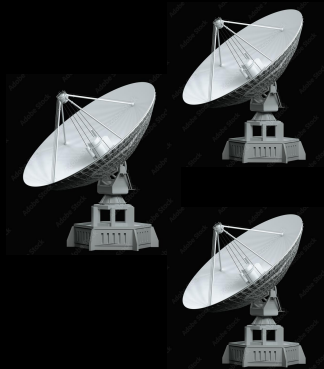
- Correlator: calculates the cross-correlation between signals received by different antennas in an array.
- Gives the "complex visibility."



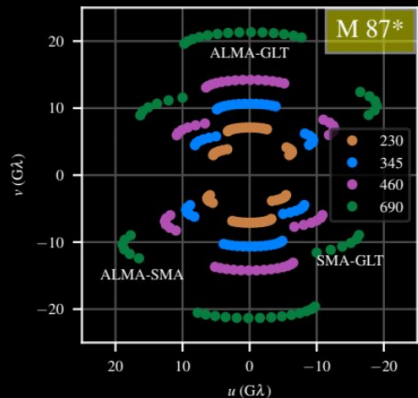
Simplified Interferometry Diagram

Van Cittert-Zernike Theorem in Interferometry

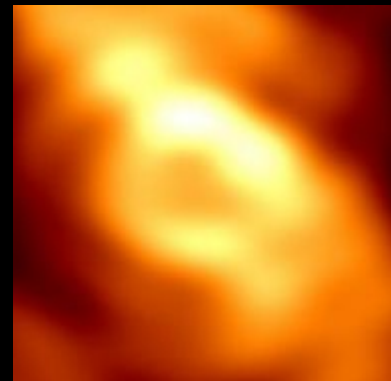
- With Fourier transform of sampled the visibilities, we can create an image.



Measured signal



Measured visibilities



FT of measured visibilities
= "dirty image"

Van Cittert-Zernike Theorem in Interferometry

- Fourier transform of sampled visibilities yields the true sky brightness convolved with the point spread function.
- Big problem: discrete points in the (u,v) plane are sampled.

$$I^D(x, y) = FT^{-1}\{B(u, v) \times V(u, v)\}$$

The diagram illustrates the Van Cittert-Zernike theorem equation with three arrows pointing from descriptive labels to the terms in the equation:

- An arrow points from the label "Dirty image" to the term $I^D(x, y)$.
- An arrow points from the label "Fourier transform of sampling function at different points in the uv plane ('dirty beam')" to the term $B(u, v)$.
- An arrow points from the label "Visibilities" to the term $V(u, v)$.

Related resources: try it yourself

```
[7]: import numpy as np
import matplotlib.pyplot as plt

def gaussian_source(size, sigma):
    x = np.linspace(-size/2, size/2, size)
    y = np.linspace(-size/2, size/2, size)
    X, Y = np.meshgrid(x, y)
    return np.exp(-(X**2 + Y**2) / (2 * sigma**2))

size = 256
sigma = 30
source = gaussian_source(size, sigma)

visibility_data = np.fft.fftshift(np.fft.fft2(source))

plt.figure(figsize=(12, 6))

plt.subplot(1, 2, 1)
plt.imshow(np.abs(visibility_data), cmap='inferno', extent=[-size/2, size/2, -size/2, size/2])
plt.title('Magnitude of Visibility Data')
plt.colorbar()

plt.show()

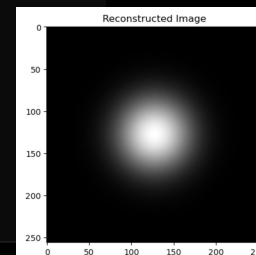
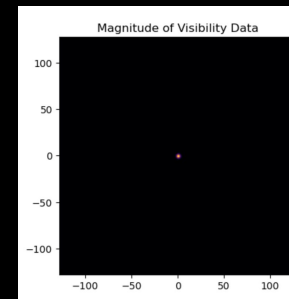
reconstructed_image = np.fft.ifft2(visibility_data)

plt.imshow(np.abs(reconstructed_image), cmap='gray')
plt.title('Reconstructed Image')
plt.colorbar()
plt.show()
```

Defining some random gaussian source

Fourier transform of source to get visibility data

Reconstruction



Related resource: Using Interferometric Data

- **Interferometric Data Simulation:** simulating visibility data to work with the Van Cittert-Zernike theorem.
 - OmniUV - ground and space VLBI observations.
 - VLBI Reconstruction Dataset - VLBI data for a variety of different telescope arrays and targets.
 - EHT Imaging
 - CASA - software used in radio astronomy for processing interferometric data (ALMA and VLA).

Useful Links!

- [Overview of the Van Cittert-Zernike theorem](#)
- [Coherence and complex signals](#)
- [Aperture synthesis explained](#)
- [Fourier transforms of images](#)
- Interferometric Data: [OmniUV](#) , [VLBI Reconstruction Dataset](#) , [EHT Imaging](#) , [CASA](#)