

Star Formation Theory-- Topics

ASTR 541
Seminar in Theoretical
Astrophysics

- **Giant Molecular Clouds (GMCs) -- structure and turbulent properties**
 - Cloud properties and observed scaling laws
 - MHD turbulence -- power spectra of v and B from simulations
 - Turbulent dissipation and driving
 - Density structure -- PDFs and correlations
 - Observational diagnostics of turbulence and magnetic fields in GMCs

- **GMCs - microphysics**
 - Heating and cooling processes for gas and grains
 - Molecule formation and dissociation
 - PDR solutions
 - Radio and IR diagnostics
 - Ambipolar diffusion, C shocks

- **Low-mass prestellar and protostellar cores: formation and structure**
 - Identification of cores in molecular-line and continuum surveys; inferring ages
 - Formation mechanisms and connection to environment
 - Bonnor-Ebert spheres
 - Magnetized equilibria; ambipolar diffusion

- **Collapse of low-mass cores**
 - Isothermal unmagnetized collapse: Larson-Penston solution, Shu solution, generalized solutions
 - Observational diagnostics of collapse
 - Effects of magnetic fields
 - Effects of rotation
 - Disk formation and magnetic braking

- **Protostars and pre-main sequence stellar evolution**
 - First hydrostatic core
 - Deuterium burning
 - Stellar birthline
 - Hayashi and Henyey tracks

- **T Tauri systems -- radiation diagnostics**

- Spectral Energy Distributions (SEDs): star, disk, and envelope contributions
- Inferred vertical structure of disks
- Accretion shock diagnostics
- Disk diagnostics from line emission

- **Winds, jets, and outflows from Young Stellar Objects (YSOs)**

- Magnetohydrodynamic (MHD) winds: disk winds, X-winds; acceleration and collimation
- Magnetosphere-disk interactions and flow onto the star
- Molecular outflows from wind/ambient interactions
- Herbig-Haro objects

- **Protostellar disks -- accretion processes**

- Thermal/ionization structure of disks; dead zones and layered accretion
- Magnetorotational instability (MRI) in partially-ionized disks -- resistivity, Hall effect, ambipolar diffusion
- Angular momentum transport via self-gravitating instabilities
- Spatially-variable accretion rates and the FU Ori outburst phenomenon
- Evolutionary models

- **Binary star formation**

- Statistics of binaries
- Formation via capture
- Formation via fragmentation of rotating core or disk
- Accretion in binary systems

- **Protostellar disks -- gas/grain dynamics**

- Regimes of grain/gas aerodynamic drag laws
- Vertical settling of dust
- Differential mass loading and Kelvin-Helmholz instabilities
- Gravitational instabilities in dust-loaded layers (Goldreich/Ward)
- Gas-dust streaming instability
- Particle concentration in gas disk structures (pressure maxima, vortices)

- **Planet formation -- overview in the star formation context**

- Planetesimal formation -- collisional agglomeration of > meter-sized bodies
- Runaway growth and oligarchic growth stages
- Gas giants -- core accretion vs. gravitational instability models
- Planet migration -- Type I and Type II
- Planet-planet scattering

- **Disk clearing around low-mass stars**

- Photoevaporation by external radiation (EUV)
- Photoevaporation by a central object (FUV, X-ray)
- Gap opening and evolution to transition disks

- **The Initial Mass Function (IMF)**

- Relation between observed IMFs and Core Mass Functions (CMFs)
- Winds/outflows and core-to-star conversion efficiency
- Theories of gravoturbulent fragmentation of GMCs into cores -- physical dependence, numerical simulations
- Competitive accretion theories

- **Star cluster formation**

- Fragmentation of massive, turbulent cores
- Feedback effects in clustered environments
- Stellar mass segregation
- Small-N dynamical interactions; sub-cluster merging
- Gas expulsion and post-expulsion dynamical relaxation

- **High-mass star formation**

- IRDCs; progenitor structures for high-mass stars
- Effects of radiation pressure on dusty envelopes/accretion flows
- Outflows from high-mass stars
- Compact HII regions

- **Destruction of GMCs**

- HII region expansion -- classical solutions
- Blister HII regions and champagne flows
- Effects of radiation pressure, stellar winds, supernovae
- Cloud lifetimes

- **Star formation on galactic scales**

- Kennicutt-Schmidt and other empirical laws
- GMC formation mechanisms -- collisional agglomeration, self-gravitating instability, effects of spiral structure
- Star formation efficiency in GMCs
- Self-regulation of star formation via feedback -- ISM equilibrium models

- **Primordial star formation**

- H₂ formation under high-z conditions
- Cooling of metal-free and metal-poor gas
- Stages of evolution for primordial prestellar "cores"
- Feedback effects in primordial star formation
- Fragmentation in primordial star formation